#### CS 6620 Shading

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# Shading models









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#### Lambert's cosine law

• Light reaching surface is proportional to projected area:  $\cos \theta$ 





### Lambertian shading

Comes from a "rough" surface (at microscopic level
Simple: light that reaches the surface is reflected equally in all directions



# Lambertian shading

 $(\overline{N}\cdot\overline{L})C_L$ 

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Color at surface:
 (where N and L are unit vectors)

# Ambient light

With this mechanism, the light in a shadowed region is 0 (black)
To avoid this, use "ambient" lighting
C<sub>ray</sub> = C<sub>surface</sub> [(N · L)C<sub>Light</sub>K<sub>d</sub> + C<sub>ambient</sub>K<sub>a</sub>]



# Guessing ambient light

It should be the average color of surfaces visible to the point being shaded
What does that imply about outdoors?

# Two-sided lighting

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- What if light hits the back of a polygon
- Options:
  - Black on back
  - Different materials for front/back
  - Two sided lighting:
    - okay:  $\overline{N} \cdot \overline{L}$  better:

Negate  $\overline{N}$  if  $\overline{N} \cdot \overline{V} > 0$ 



# Two-sided lighting

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Ray2

- You might consider checking if the light is on the "right side" of the object
  - $\operatorname{sign}(\overline{-V} \cdot \overline{N}) = \operatorname{sign}(\overline{L} \cdot \overline{N}) \operatorname{Ray1}$  $(\overline{V} \cdot \overline{N})(\overline{L} \cdot \overline{N}) < 0: \operatorname{lit}$  $(\overline{L} \cdot \overline{N}) > 0: \operatorname{lit} \text{ (when normal flipped)}$
- Can avoid casting shadow rays too
   Or use absolute value for L·N
  - Or use absolute val

# Two-sided lighting



Left: Normal two-sided lighting (Abs L·N)
Right: Comparing signs

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#### Lambertian Shading

Compute hit position  $(\vec{P} = \vec{O} + t\vec{V})$ Call primitive to get normal  $(\overline{N})$  (normalized)  $costheta = \overline{N} \cdot \overline{V}$ if(costheta < 0)normal =-normal Color light = scene.ambient\*Ka foreach light source get  $C_{L}$  and  $\overline{L}$ dist= $\|\overline{L}\|, \overline{L_n} = \frac{\overline{L}}{\|\overline{L}\|}$  $cosphi = \overline{N} \cdot \overline{L_n}$ if(cosphi > 0)if (! intersect with 0 < t < dist) light  $+= C_1 * (Kd * cosphi)$ result=light\*surface color

# Background

- What do you do when the ray misses all objects?
  - Constant color
    - Assign color to ray
  - Gradient
    - Map -1 to 1:
    - Map to colors:
  - Star field:
    - Sum up colors:
  - Environment map

$$= \frac{V_{ray} \cdot V_{up}}{\left\|V_{ray}\right\| \left\|V_{up}\right\|}$$

$$C_{down} + \left(\frac{s+1}{2}\right)C_{up}$$

$$\sum_{stars} \left(\overline{V_{ray}} \cdot \overline{V_{star}}\right)^{p} C_{star}, p \approx 10000$$



# Light transport

There are 4 primary ways that light interacts with a surface (are there hybrids?):

- 1. Bounces off (perfect specular reflection)
- 2. Absorbed and retransmitted in an arbitrary direction (perfect diffuse reflection)
- **3**. Travels through surface (perfect specular transmission)
- 4. Absorbed and retransmitted on other side (perfect diffuse transmission)



#### Reflected and transmitted rays

plane of incidence



boundary

# Reflections

- Reflection can be computed by tracing another ray from the intersection point
   Called perfect approximation
- Called perfect specular reflection





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# Reflections

Perfect "bounce": incoming angle equal to outgoing angle •  $\theta_{I}$  is called "angle of incidence" or "incident angle" • V, N and P define a plane Trace new ray on the same plane with origin at P and direction R Two derivations: algebraic and geometric



#### Geometric derivation

S

V

Ν

 $\theta_{I=}\theta_{r}$ 

R

S

N cos  $\theta_i$ 

 $\theta_{\rm r}$ 

 $\overline{R} = 2\overline{N}\cos\theta_i + \overline{V}$  $\overline{R} = 2\overline{N}\left(-\overline{N}\cdot\overline{V}\right) + \overline{V}$  $\overline{R} = \overline{V} - 2\left(\overline{N}\cdot\overline{V}\right)\overline{N}$ 

#### Reflected color

In Material::shade, trace and shade a ray to get the color along R
Reflectivity r (range 0-1)
Result color = r \* color along ray
More about the reflectivity later

#### Transparency

Light is an electromagnetic wave
 Some materials conduct energy at visible frequencies







# Transmitted rays

plane of incidence



boundary

#### Maxwell's equations

 Maxwell's equations relate electricity and magnetism (light) as waves in space

They predict the speed of light in a vacuum:

 $c = \overline{\sqrt{\mu_0 \varepsilon_0}}$   $\mu_0$ : Electric permittivity of free space  $\varepsilon_0$ : Magnetic permeability of free space

# Speed of light

Speed of light changes (slows) in other materials
 v = 1/õε
 µ : Electrical permittivity of material
 ε: Magnetic permeability of material
 v ≤ c



#### Fermat's principle (1657)

#### Light takes the fastest path between two points: $t_{AB} = t_{AC} + t_{BC} = \frac{\|C - A\|}{v_a} + \frac{\|B - C\|}{v_b}$ A

Simplify:  $B_v = 0$ 

$$t_{AB} = \frac{\sqrt{(C_x - A_x)^2 + A_y^2}}{v_a} + \frac{\sqrt{(B_x - C_x)^2 + B_y^2}}{v_b} - \frac{V_b}{v_b}$$



#### Refraction





#### Index of refraction

Absolute index of refraction:

 $\eta_{abs} = \frac{c}{v}$ 

where  $c = 2.99 \times 10^8$  m/s. Note:  $v \le c$  for all transparent materials, so  $\eta \ge 1$ .

Relative index of refraction:

$$\eta = \frac{\eta_2}{\eta_1}$$



#### Snell's Law

#### plane of incidence



 $\frac{\sin \theta_1}{\sin \theta_2} = \frac{v_1}{v_2} = \frac{\eta_2}{\eta_1} = \eta = \eta_{12}$ 

boundary



# Questions?